

Automatic Generation of Caustic Maps to Enhance Microlensing Modeling of Lensed Quasars light curves

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Aug 2023



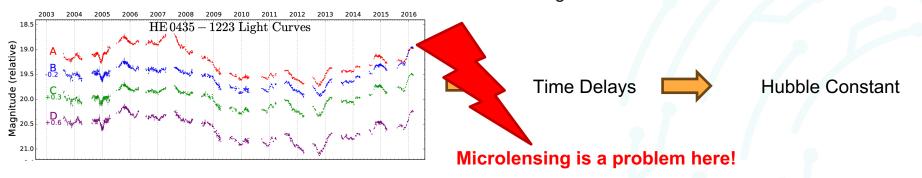


Modeling Microlensing in light curves of Lensed Quasars

Multiply-lensed Quasars

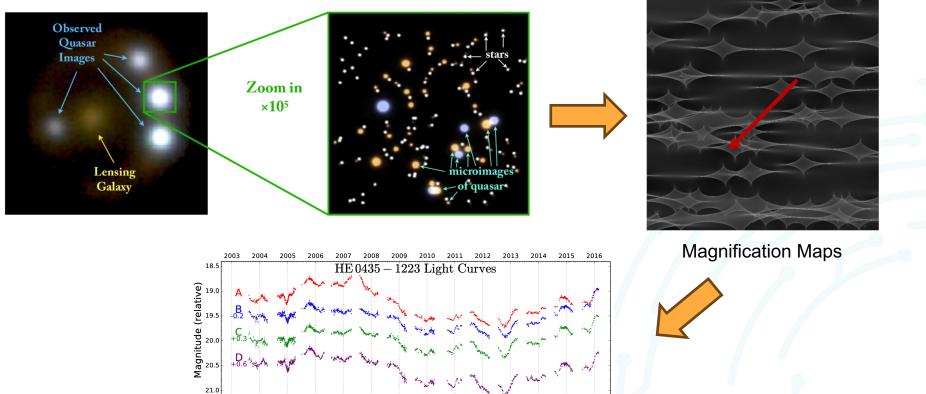


- The multiple images will be affected by the individual stars and objects in the lens galaxy along their line of sight, which is called "microlensing".
- Enhanced modeling of the microlensing variation in lightcurves of strong-lensed multiply-imaged quasars leads to more precise measurements of the cosmological time delays and the Hubble Constant.
- We can use it not only to get better measurements of H₀, but also to study quasars structure and the distribution of mass of stars in distant galaxies.





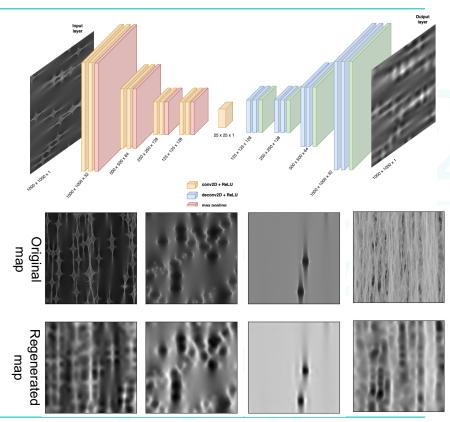
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Caustics Maps Generation using Al

- We use variational autoencoders (VAE) to compress magnification maps to speed up their generation or to even bypass maps altogether.
- Our training set includes ~12000 pre-computed magnification maps (Vernardos et. Al. 2014).
- The structure of our VAE is inspired by the Unet² network that is used for image segmentation.
- The network is able to reduce the dimensions from 1000x1000 to 25x25 and back to the original dimensions and regenerate the patterns in the maps.
- Different parameters can generate maps that have significantly different patterns and statistical features. The VAE is able to capture all these differences.





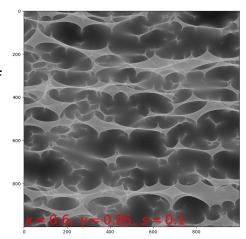
Results

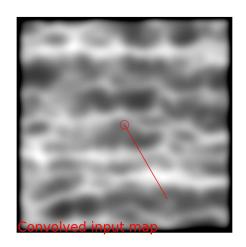
 The latent space representation is explored in 2D visualizations to see how maps with different ranges of initial parameters are distinguished. The reduced dimensions form a small meaningful parameter space that contain the most important features in the input.

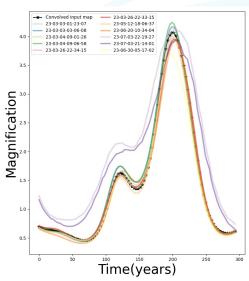
• We generate light curves from the predicted maps and compare with the true light curves to find how

details of the maps are reproduced.

 Next step is to make this dimension reduction part of the lightcurve analysis and help automate the process.









Thanks for listening!

