### Search for anomalous objects in the ALLWISE survey

#### Agnieszka Pollo

National Centre for Nuclear Research (Warsaw) & Jagiellonian University (Cracow); Poland with Aleksandra (Ola) Solarz; Maciek Bilicki, Mariusz Gromadzki, M., Anna Durkalec, Michal Wypych

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Based on the paper: "Automated novelty detection in the WISE survey with one-class support vector machines", Solarz et al. 2017 Astronomy & Astrophysics, Volume 606, id.A39

# Wide-field Infrared Survey Explorer (WISE)

#### Solarz et al. 2017

- All-Sky survey in IR
- Detected over 747 mln sources

(15 PB of data; tables + images)

- Publicly available (position, photometry in 4 bands (3.6-22 um))
- Low angular resolution (~6")
- No redshift information



(http://wise2.ipac.caltech.edu/docs/release/allsky/)

## Exploration of parameter spaces

→ The usual first approach to selection: CC diagrams

→ Novel or anomalous sources deviate from expectations but could mimic a behaviour of a 'normal' object

→ need for new ML approaches based on larger parameter spaces



Credit: Wright+10

### WISE: first step towards ML novel source detection

Training set:

 $\rightarrow$  AllWISE x SDSS ( $\alpha$ , $\delta$ ) with secure spectro-z



#### Solarz et al. 2017



### WISE: first step towards ML novel source detection

#### Solarz et al. 2017





Parameter space: → Brightness W1 → Color W1-W2 → Compactness W1mag13 = w1mpro(5") - w1mpro(11")

### WISE: accounting for unknown Solarz et al. 2017 Unknowns



## Novelty detection with One-Class Support Vector Machines



- Create one 'known' class (mix of AllWISE x SDSS galaxies, stars, QSOs)
- Maps input data to a higher D parameter space (based on Kernel methods)
- Hypersurface hugging the expected sources
- Anything with 'unknown' patterns falls outside the hypersurface => novelties





~650,000 anomalous sources

Solarz et al. 2017

What are they?

## **Spurious sources**

- W1-W2 ~ -1 ; 80%; Spitzer GLIMPSE: IRAC I1 [3.6 um], IRAC I2 [4.5 um]
- Low WISE resolution (6") in crowded fields => blends
- OCSVM: good tool for selecting hidden artefacts





Solarz et al. 2017



## AGN candidates?

- 30,000 sources (those in Galactic Plane: still mostly blends)
- 76% undetected at other wavelengths!
- ~7 000 objects with SDSS photometry (no spectro-z), follow-up observations
  - Peculiar (dusty) QSOs
  - Low-z very dusty galaxies
  - Very dusty Galactic objects like YSO







## Conclusions

- With sufficiently big data (like LSST) anomalies may come in big numbers
- Parameter space is important
- What we find is mostly trash but
  - valuable trash if not found by other methods  $\rightarrow$  novelty search as a method for additional cleaning of catalogs
  - some genuine atypical sources
  - nedeed methods of separation (clustering in different parameter spaces/with DL)