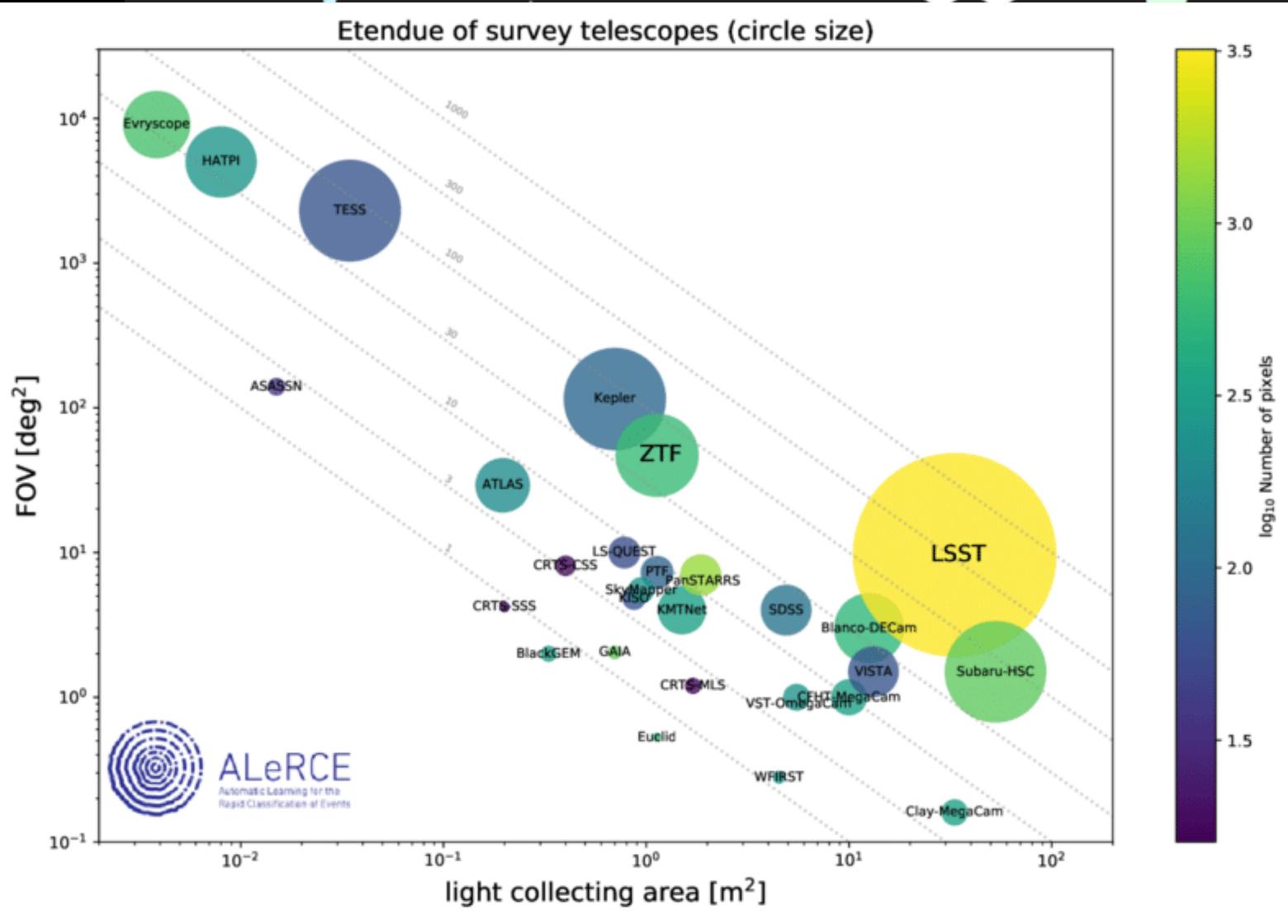


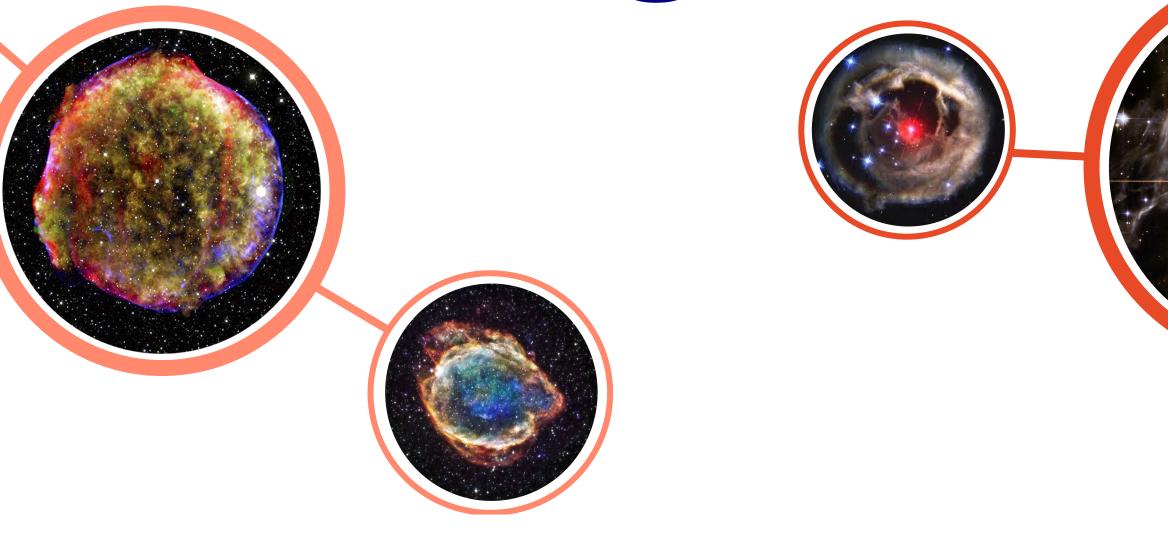
# ETENDUE ENABLES TIME-DOMAIN ASTROPHYSICS

- With a 8.4m primary mirror...
- and a unique f/1.23 optical design giving it a very large FoV...
- Rubin Obs. has the largest etendue of any groundbased optical telescope
- leading to ~10 million transient alerts, every night

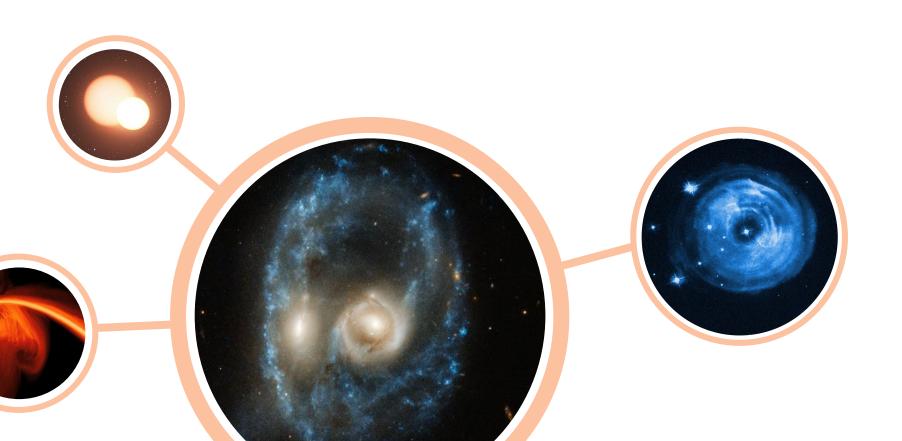


The Big Picture

- Understand the physics of stellar death
- probe transient demographics
- discover the most rare and faint sources
- map the covariance between transients and their host environment, and their evolution with redshift
- study the Hubble constant tension, dark energy and its evolution over over cosmic time

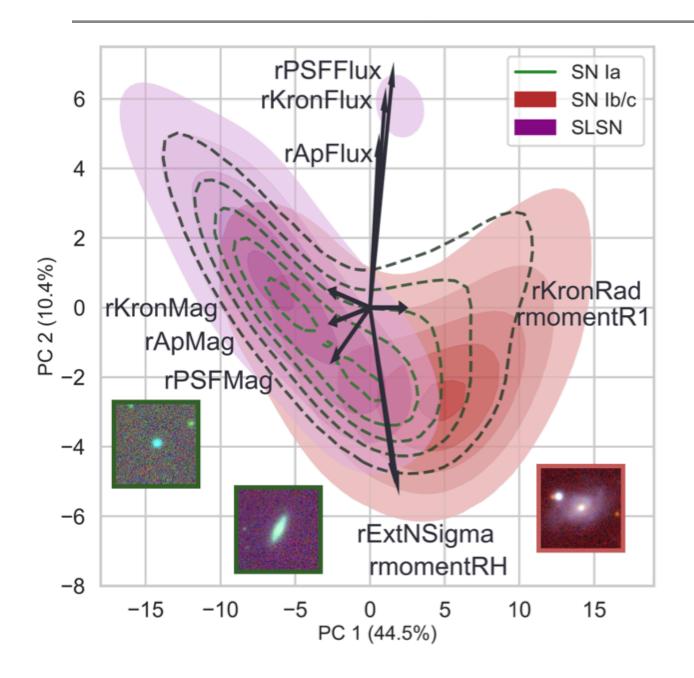


# LSST ALLOWS US TO EXAMINE SEVERAL OPEN QUESTIONS



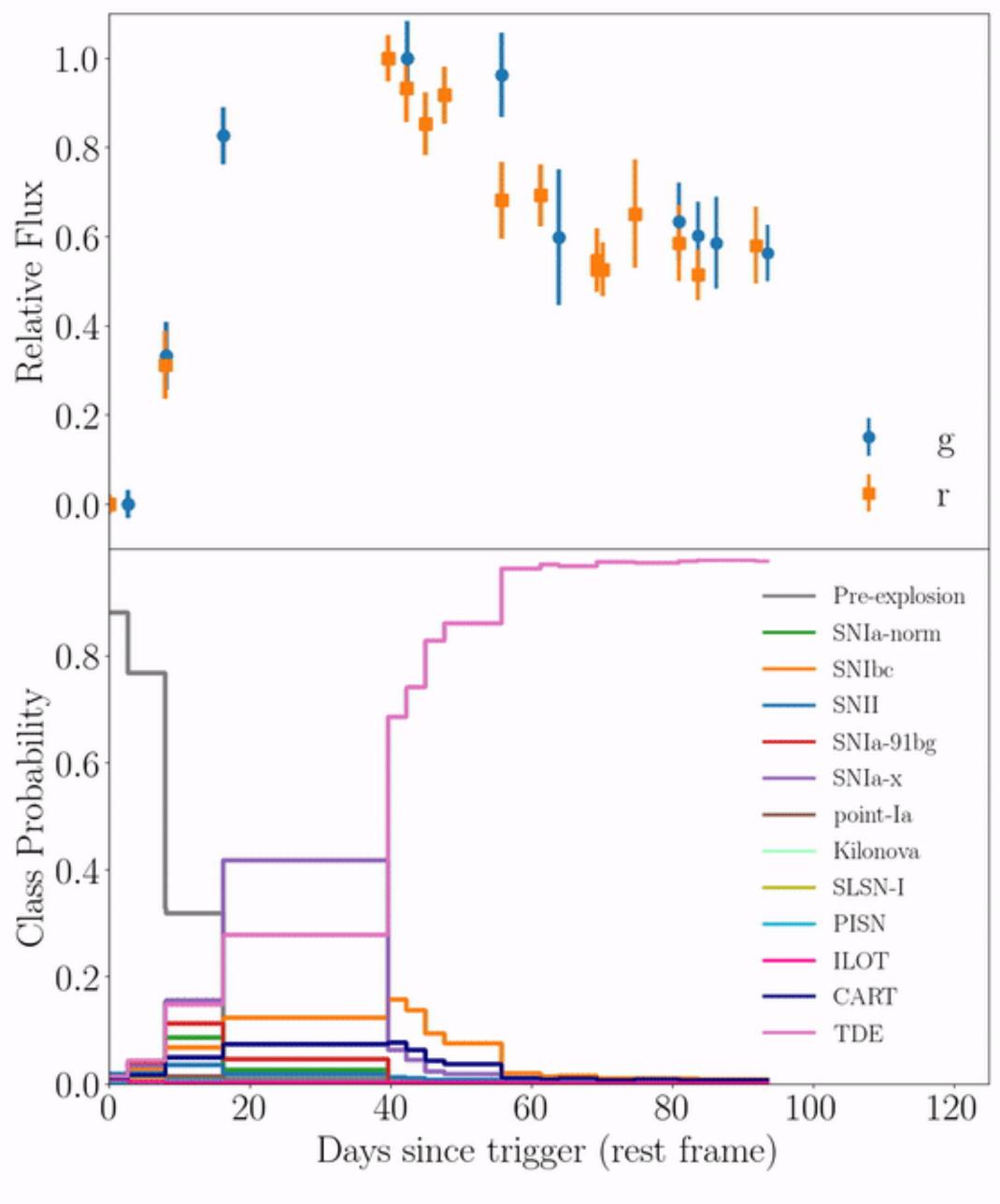


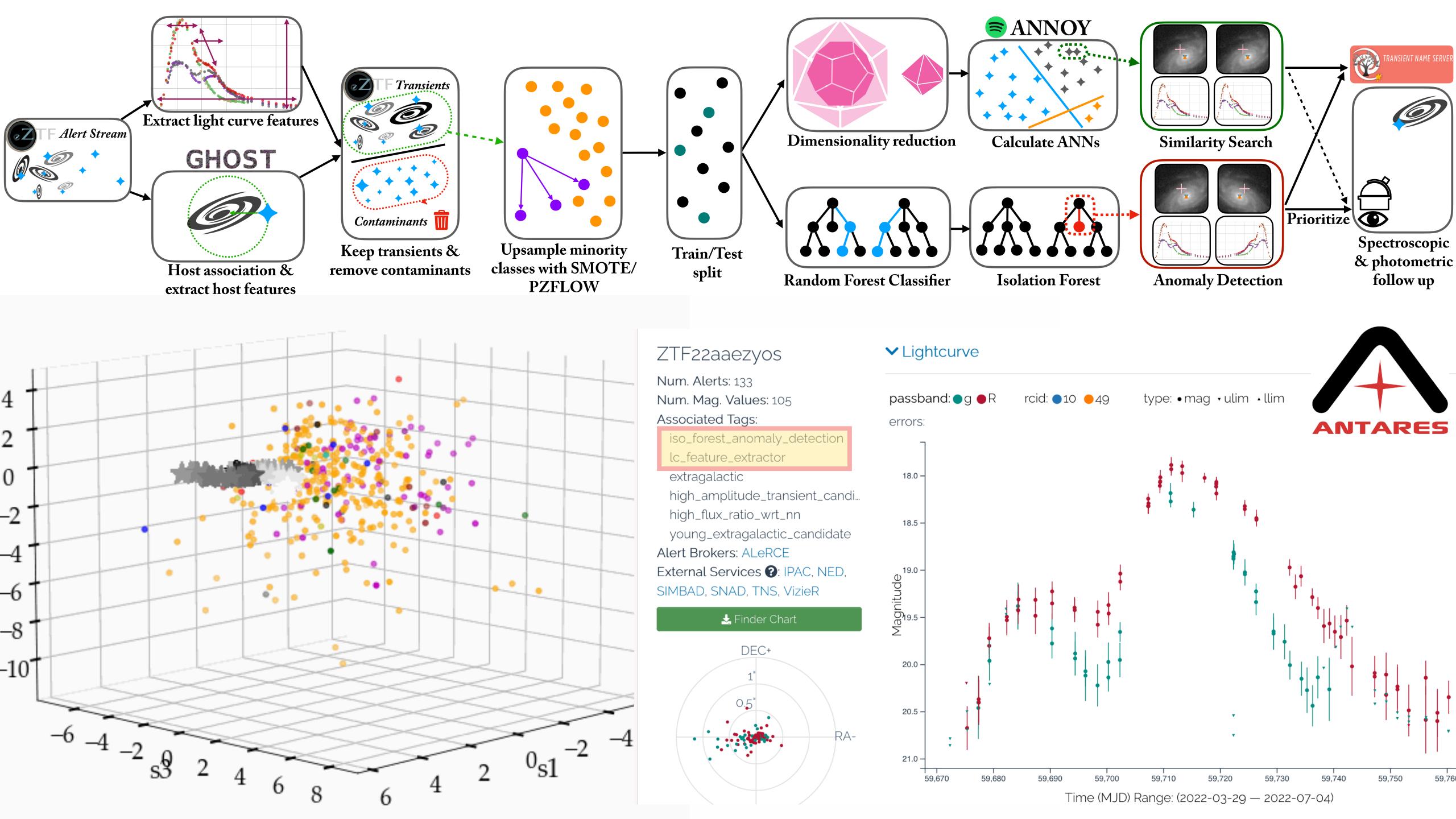
### THE ROLE OF AI WITH LSST

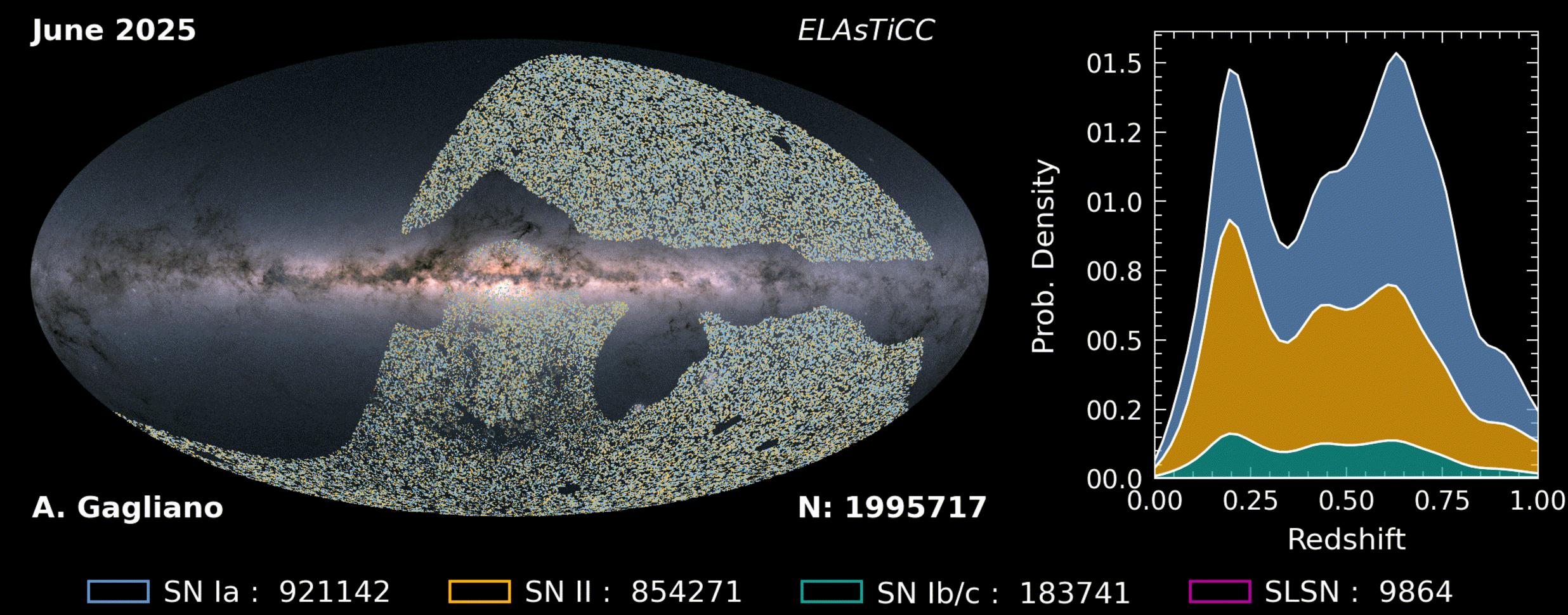


RAPID (Right, Muthukrishna, Narayan+19) uses real-time alerts to classify objects, while GHOST (left, Gagliano, Narayan+21) uses only host-information to classify SNe

- The key challenge is how to enable a community with diverse interests to use the vast quantity of data
- For the time-domain, need to process real-time streams to characterize events quickly implies AI/ML
- The primary goal is to whittle 10 million/alerts per night into something manageable for your science case







- Preparing AI for LSST requires detailed simulations of the time-domain sky beyond DP2 together with connecting with machine learning-based "brokers" such as ANTARES (Narayan+'18, Matheson+'21)
- Current best realization of these simulations is the Extended LSST Astronomical Time-series Classification Challenge (ELAsTiCC) Round 2 starts at the end of August!
- December 2023 NASA effort (Leads: Michael Troxel, Rachel Mandelbaum, Mike Jarvis) to create pixel level simulations of Rubin+Roman sky

## ELAsTiCC has a rich diversity of models

- New:
  - delta Scuti, Cepheids (K. Malanchev)
  - dwarf novae and changing look AGN (Q. Cheng)
  - + others we've snuck into test data but not in training data
- Updated:
  - more diverse SNe Ib & c (M. Vincenzi)
  - M-dwarf flares (V. Shah)
  - ◆ KNe (V. Shah, D. Chatterjee)
    - + synthetic LVK O4 alert skymaps

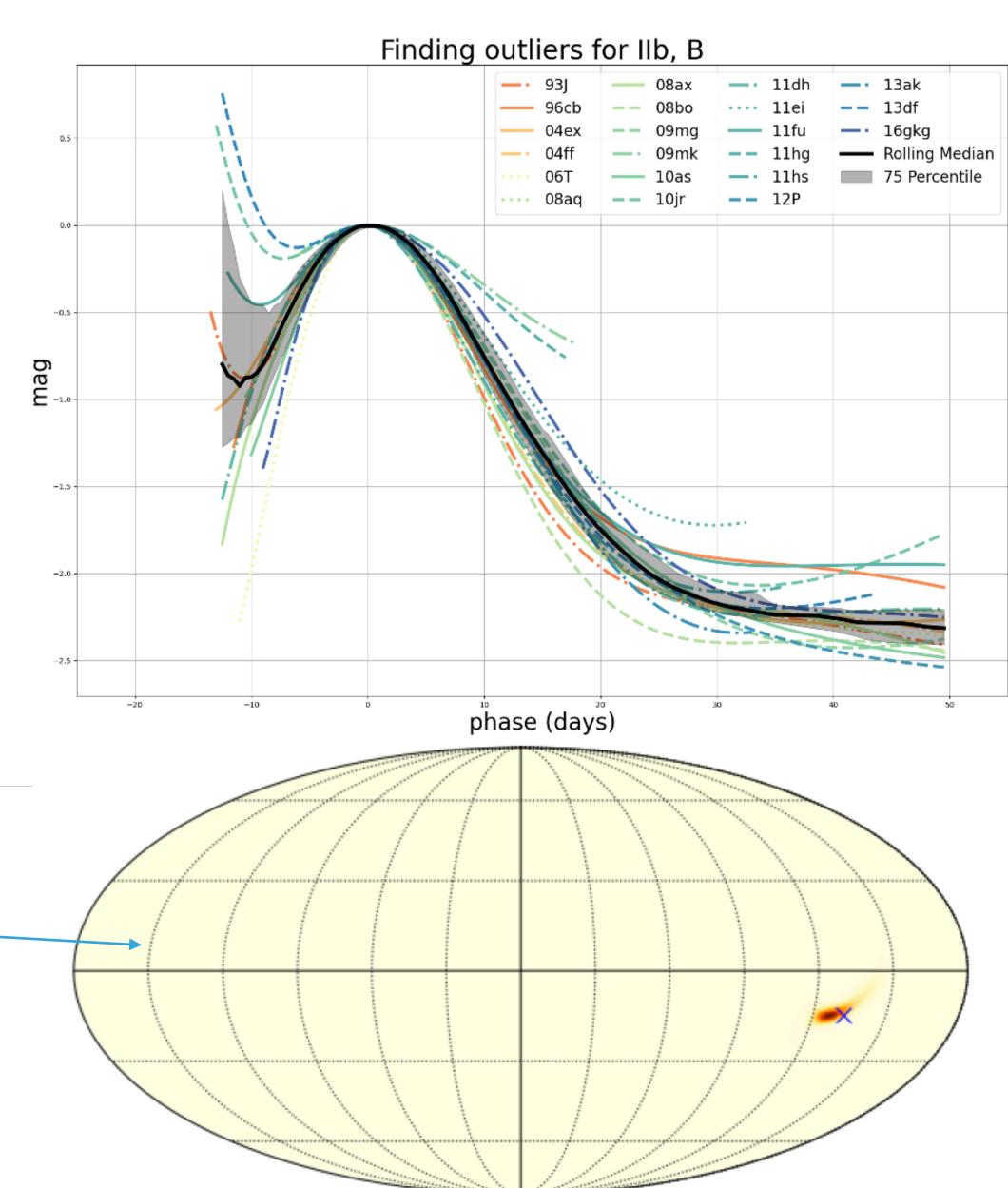












### What is ELAsTiCC?

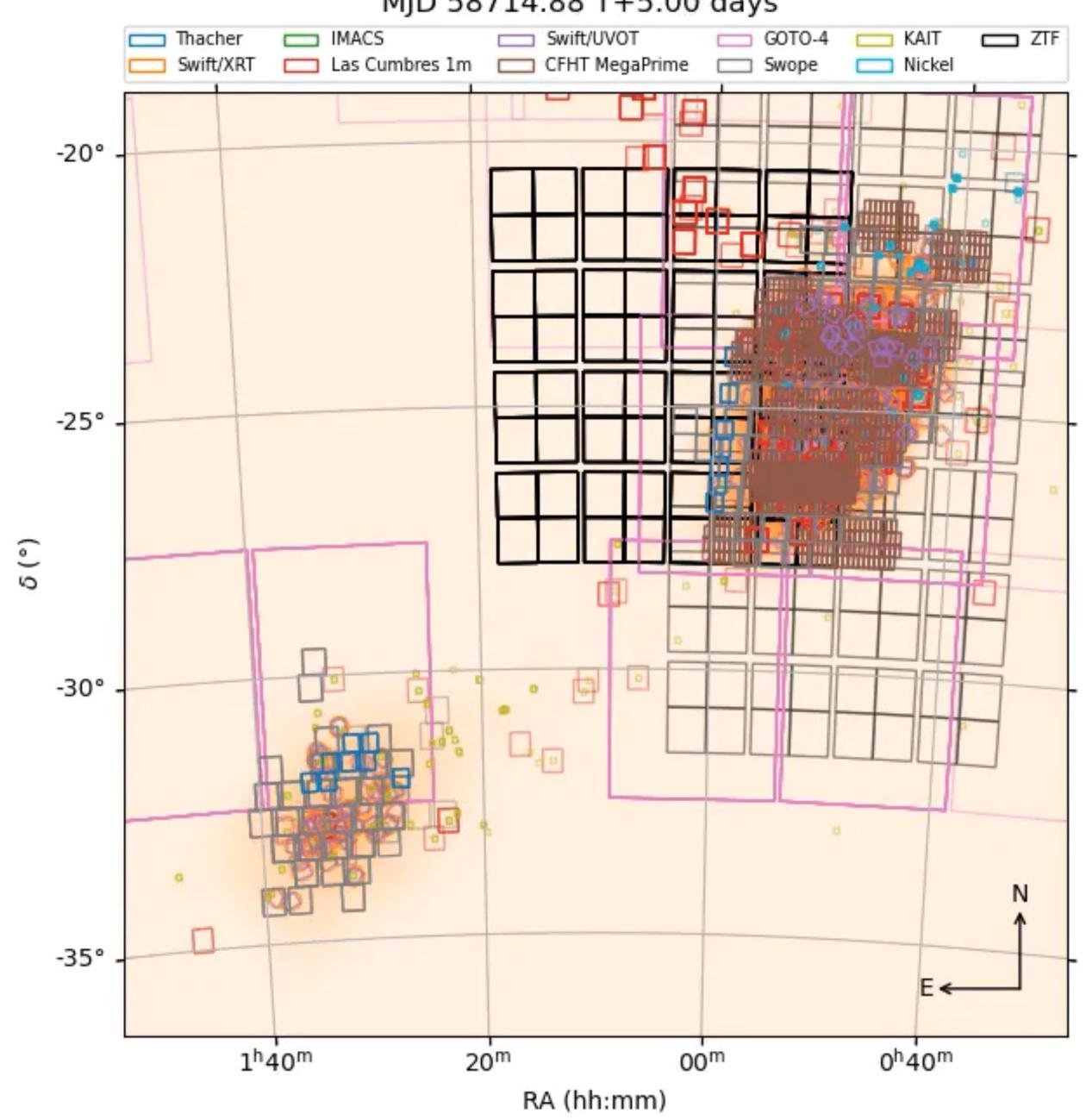
Big infrastructure challenge with multiple potential points of failure along chain. ELAsTiCC simulates the entire chain to identify these.



https://portal.nersc.gov/cfs/lsst/DESC\_TD\_PUBLIC/ELASTICC/

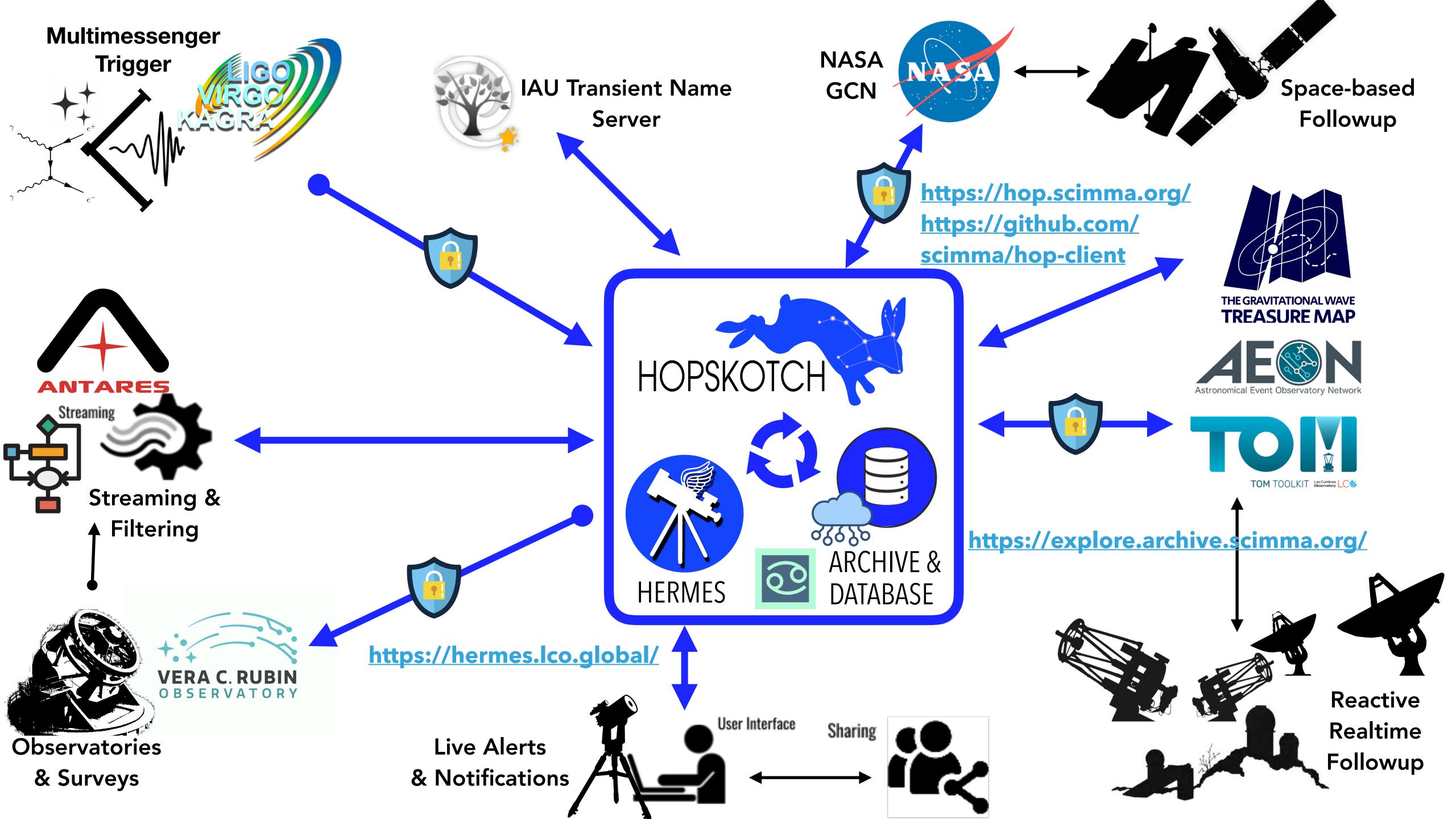
#### GW190814 Followup

MJD 58714.88 T+5.00 days



- LSST does not operate in isolation!
  - Critical to coordinate with other optical surveys such as the Young Supernova Experiment (YSE), Roman, Euclid, DESI2 as well as MMA experiments - LVK, IceCube, CTA, CMB-S4 to identify rare, faint and fast sources
- ▶ This needs an interconnected ecosystem for time-domain astrophysics
- ▶ This is the goal of the NSF CSSI-funded Scalable Infrastructure for Multi-Messenger Astrophysics team (SCiMMA)
  - Need new features? Let's talk!







- ▶ DESC and SCiMMA are building this infrastructure and the datasets to prepare for LSST
- We're happy to work with you and share code/add new models/features! Let's talk!