

Objectives & Research Question:

- Characterize SN hosts using largest sample to date (ZTF-based).
- Test if host features can help improve classification of SN.

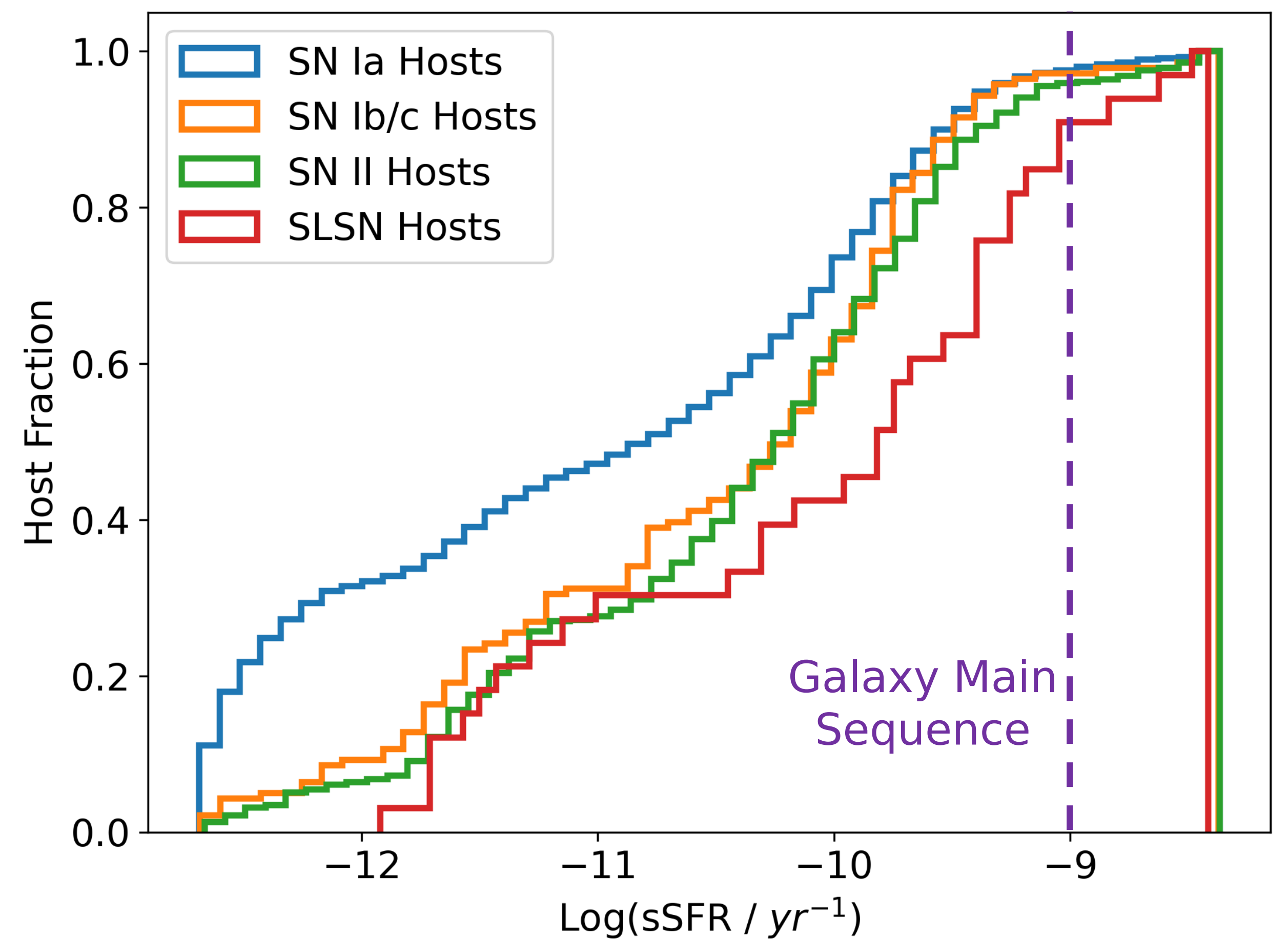
Background:

- Host galaxy properties give hints of the population of future SNe progenitors.
- SNe host galaxy properties have proven useful in classifying SNe, despite lacking information of the event itself.
- The ALERCE broker (Förster+21) has developed a successful LC classifier (Sanchez-Saez+21) using ZTF light curves. However, better distinction among SNe subtypes is needed.

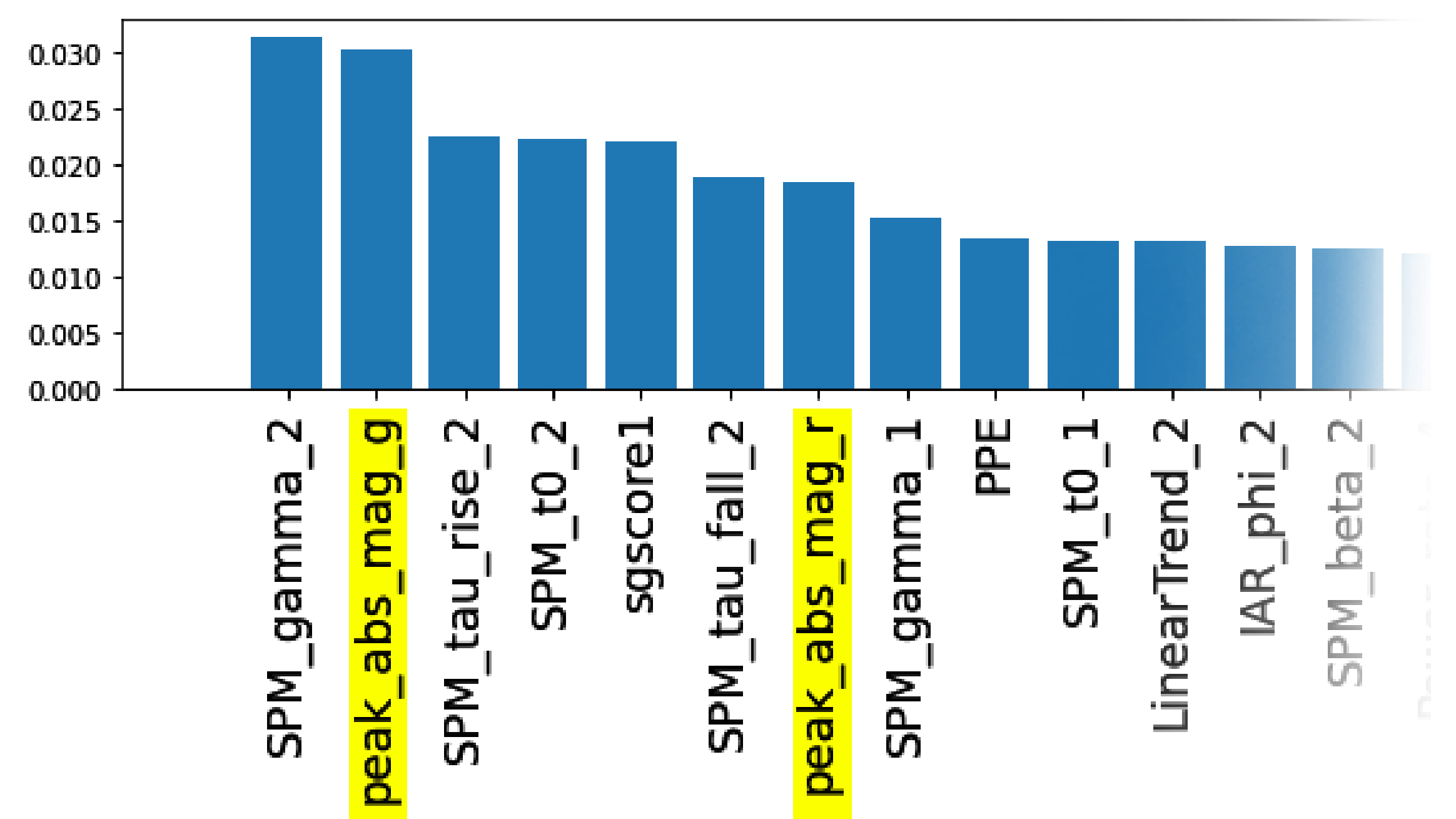
True label	SNla	SNlbc	SNII	SLSN	QSO	AGN	Blazar	YSO	CV/Nova	LPV	E	DSCT	RRL	CEP	Periodic-Other
SNla	76 ⁺⁷ ₋₆	18 ⁺⁵ ₋₆	5 ⁺² ₋₂	1 ⁺¹ ₋₁	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀
SNlbc	33 ⁺¹¹ ₋₁₁	50 ⁺¹⁷ ₋₆	11 ⁺¹¹ ₋₆	6 ⁺¹¹ ₋₆	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀
SNII	15 ⁺⁸ ₋₄	16 ⁺³ ₋₀	53 ⁺⁷ ₋₇	17 ⁺⁹ ₋₉	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀
SLSN	0 ⁺²⁸ ₋₀	0 ⁺¹ ₋₀	0 ⁺²⁵ ₋₀	100 ⁺⁰ ₋₂₆	0 ⁺⁰ ₋₀	0 ⁺¹ ₋₀	0 ⁺²⁵ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀
QSO	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	87 ⁺¹ ₋₁	8 ⁺⁰ ₋₀	5 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀
AGN	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	9 ⁺² ₋₁	85 ⁺² ₋₂	5 ⁺² ₋₁	0 ⁺¹ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀
Blazar	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	17 ⁺⁴ ₋₄	7 ⁺¹ ₋₂	74 ⁺⁵ ₋₃	1 ⁺¹ ₋₁	1 ⁺¹ ₋₁	0 ⁺¹ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺¹ ₋₀
YSO	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺¹ ₋₀	1 ⁺⁰ ₋₁	78 ⁺³ ₋₃	0 ⁺¹ ₋₀	4 ⁺¹ ₋₁	1 ⁺¹ ₋₀	0 ⁺¹ ₋₀	1 ⁺² ₋₀	3 ⁺¹ ₋₁	11 ⁺² ₋₁
CV/Nova	4 ⁺³ ₋₃	2 ⁺¹ ₋₁	1 ⁺¹ ₋₀	0 ⁺⁰ ₋₀	0 ⁺¹ ₋₀	0 ⁺⁰ ₋₀	1 ⁺¹ ₋₁	1 ⁺² ₋₁	68 ⁺³ ₋₃	1 ⁺¹ ₋₁	3 ⁺¹ ₋₁	4 ⁺² ₋₂	10 ⁺³ ₋₃	2 ⁺¹ ₋₁	2 ⁺² ₋₁
LPV	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	1 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	98 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	1 ⁺⁰ ₋₀	1 ⁺⁰ ₋₀
E	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	1 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	74 ⁺⁰ ₋₂	7 ⁺⁰ ₋₀	3 ⁺⁰ ₋₀	4 ⁺⁰ ₋₀	12 ⁺¹ ₋₁
DSCT	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	5 ⁺² ₋₂	89 ⁺³ ₋₆	2 ⁺¹ ₋₁	1 ⁺¹ ₋₁	3 ⁺¹ ₋₁
RRL	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	4 ⁺¹ ₋₀	4 ⁺⁰ ₋₁	87 ⁺⁰ ₋₁	2 ⁺⁰ ₋₀	2 ⁺¹ ₋₀
CEP	0 ⁺⁰ ₋₀	0 ⁺¹ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺¹ ₋₁	1 ⁺¹ ₋₀	0 ⁺⁰ ₋₀	0 ⁺² ₋₀	3 ⁺¹ ₋₁	1 ⁺¹ ₋₁	11 ⁺⁵ ₋₃	76 ⁺⁴ ₋₃	9 ⁺³ ₋₃
Periodic-Other	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	0 ⁺⁰ ₋₀	1 ⁺¹ ₋₁	1 ⁺¹ ₋₁	2 ⁺⁰ ₋₀	11 ⁺³ ₋₃	3 ⁺¹ ₋₁	5 ⁺² ₋₂	5 ⁺² ₋₂	73 ⁺³ ₋₃

Preliminary Results:

- Properties derived from SED fits are obtained for all SN hosts.
- Cumulative histograms are made to characterize the distribution of host properties by SN class.



- We ingested this information into a modified LC classifier finding a modest improvement largely due to peak luminosity.



Methodology:

- Select all 5283 spectroscopically classified SNe reported to TNS.
- Associate SN with their host using DELIGHT (Förster, in prep.) and pull their photometry from the Legacy Survey DR9 (Dey+19).
- Find host redshift in the literature (SDSS, PRLS) or using photometry (EAZY, Bagpipes).
- Perform z-fixed SED fitting to derive physical properties from the host (e.g., stellar mass, SFR).
- Compare with previous studies (such as Schulze+21; 888 PTF SNe).
- Add new features to the LC classifier and compare with prior performance.
- Eventually apply to all LC-classified ZTF SNe (6507) to study host statistics.

Discussion and Future Work:

- Results are as reliable as the photometry, which is poor in some cases.
- Derived physical properties and relations are broadly consistent with previous works.
- Bigger sample will provide better statistics versus previous studies.
- Peak absolute magnitudes of SNe had high levels of importance, leading to some-what (~5%) improved classification accuracy, while surprisingly other host physical properties did not. Such information about the host might be contained in some LC Classifier training features (e.g. colors).
- In the future, we aim to obtain better photometric redshifts and perform data augmentation to mitigate issues related to an unbalanced training set.

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