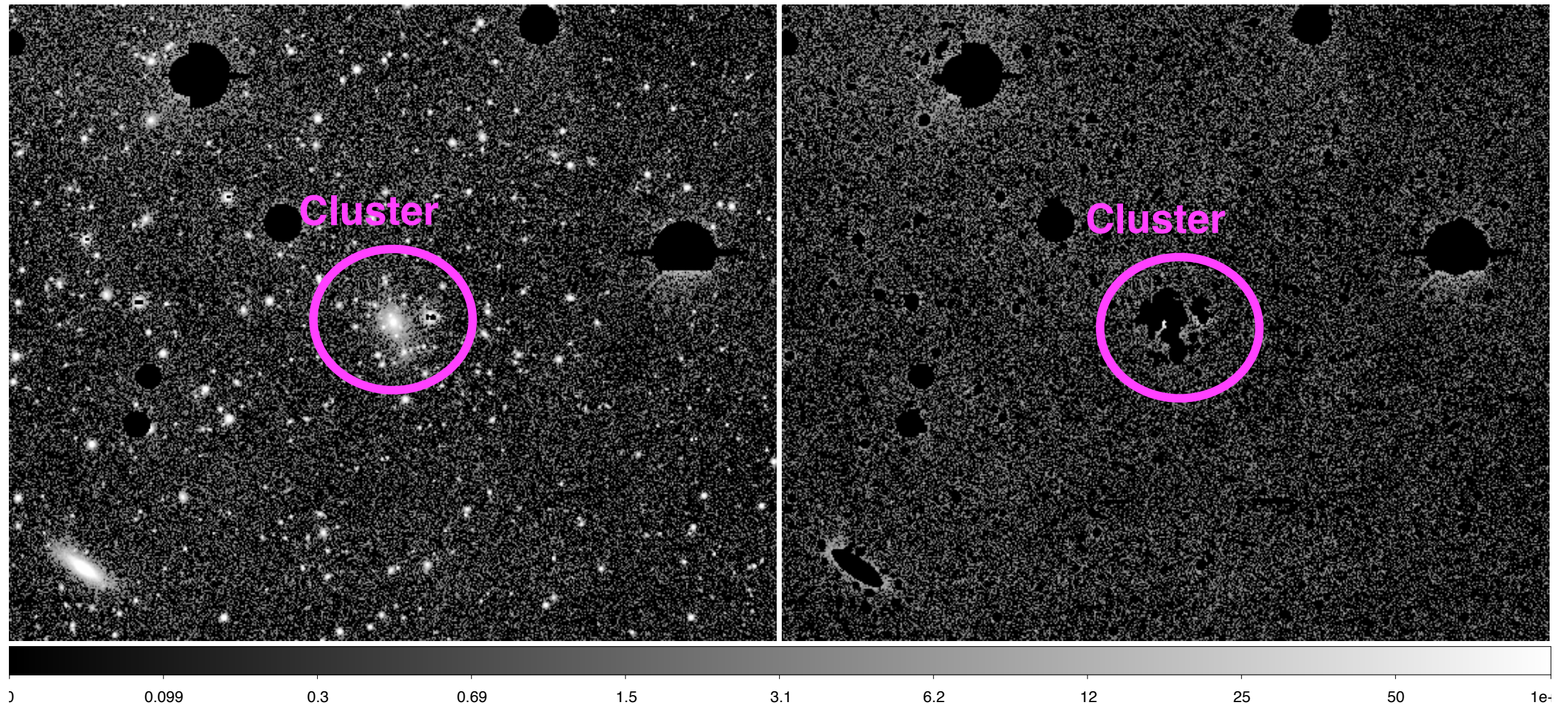


Studies of intra-cluster light through stacking Dark Energy Survey Data

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Step 1: Individual cluster images are co-added, background subtracted, and masked to detect ICL.

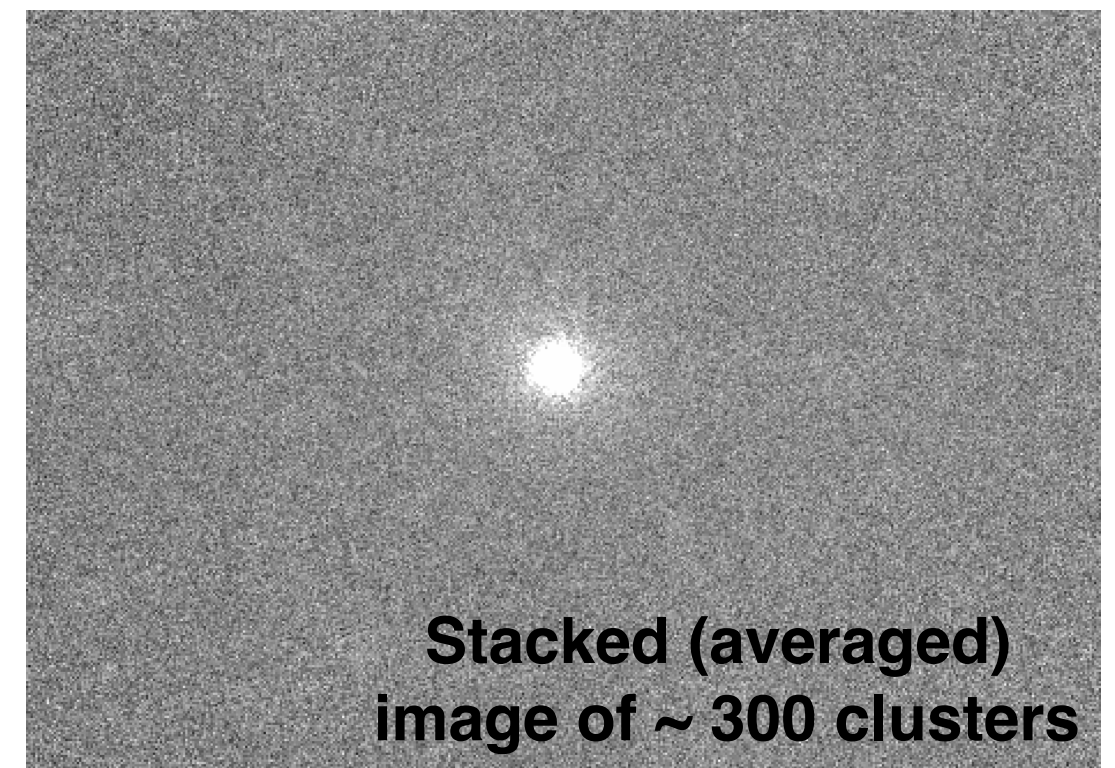
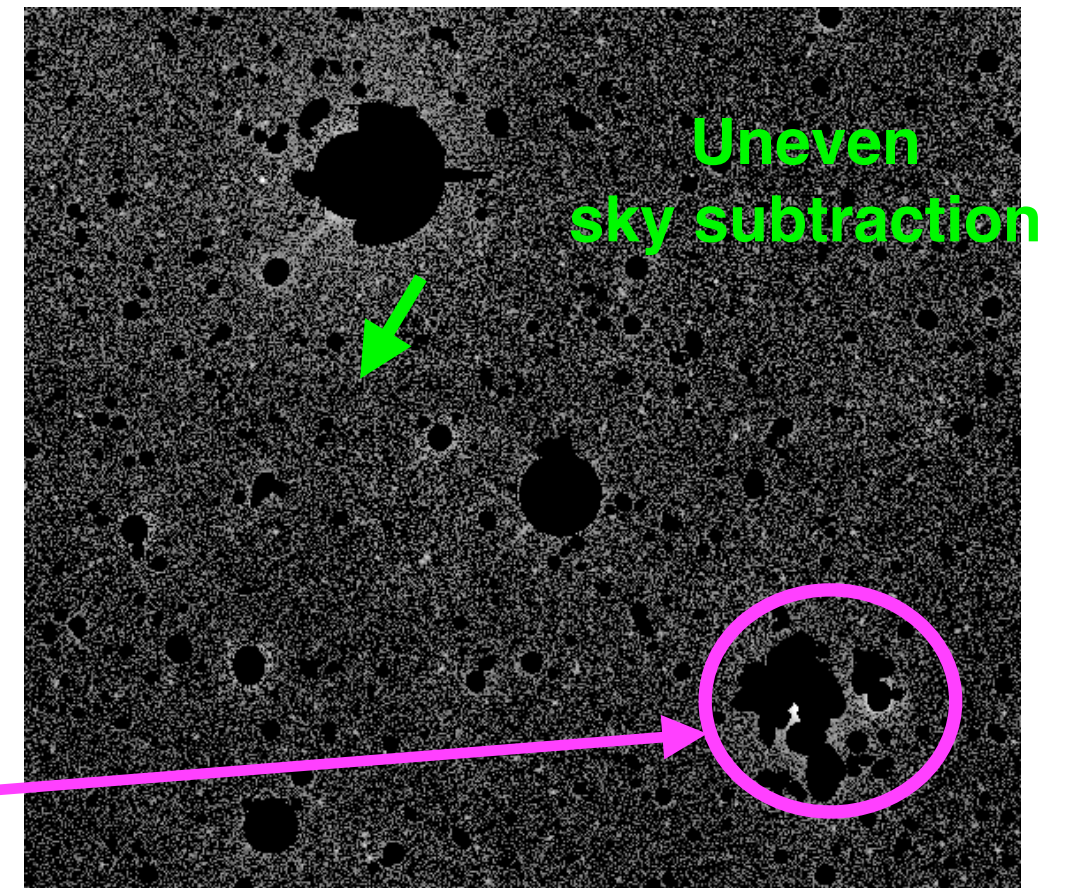
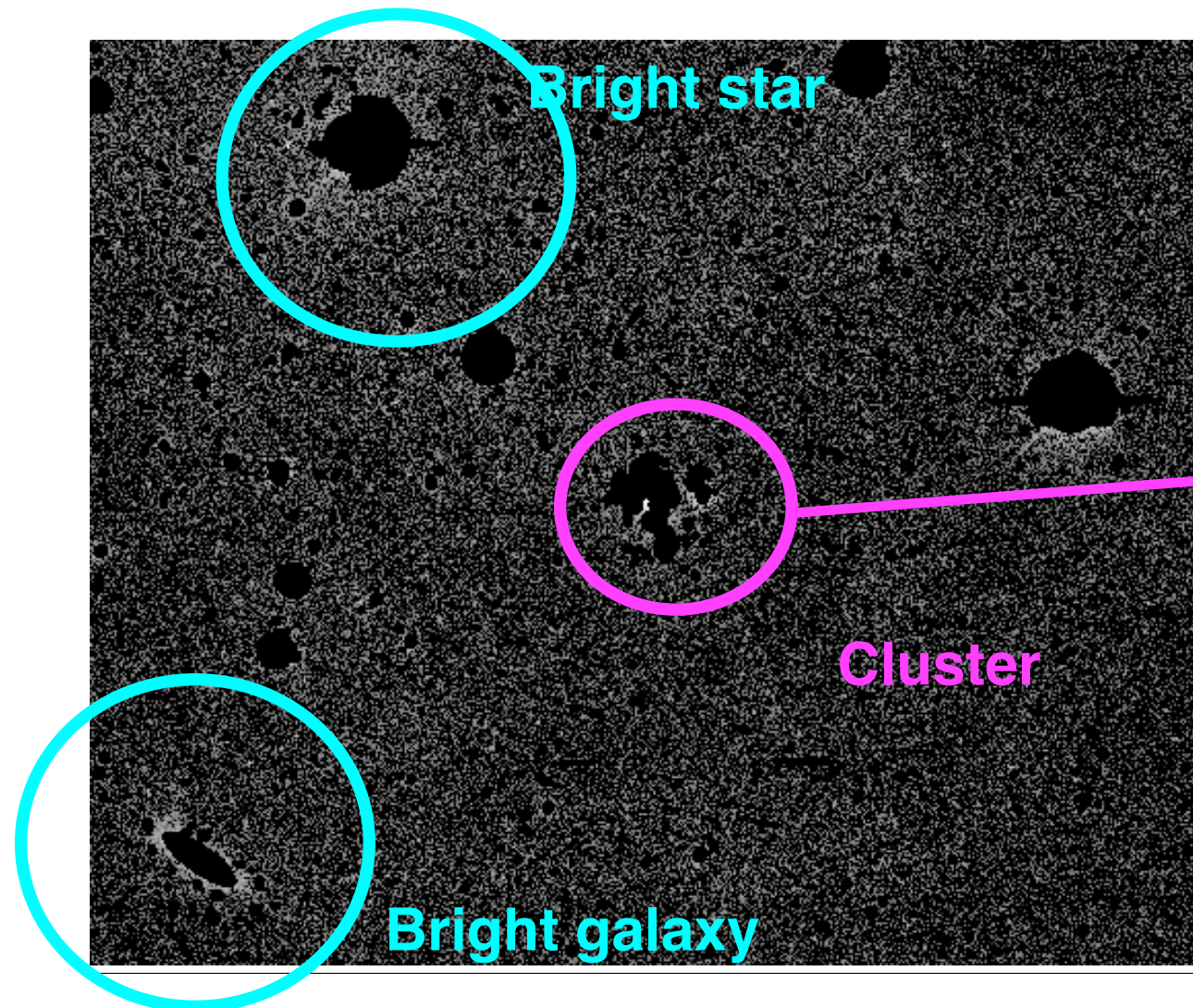


Based on DES Y3 data

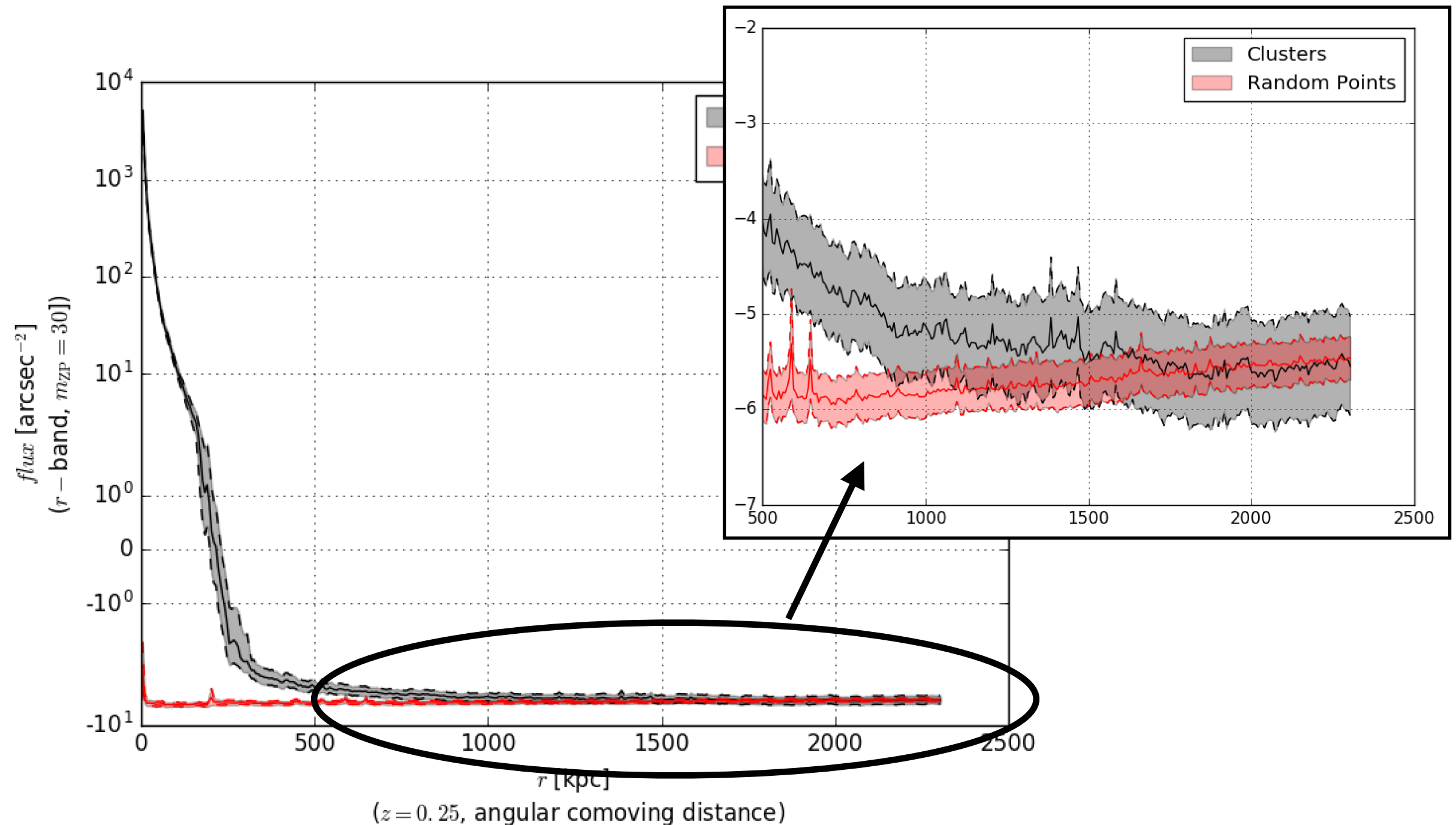
Image of one cluster,
of redshift 0.23, richness 57

Sky background is estimated over the whole FOV, $\sim 3 \text{ deg}^2$
(Bernstein+2017).

Step 1: Many effects still persist, but “stacking/averaging” helps removing some of the un-correlated effects.

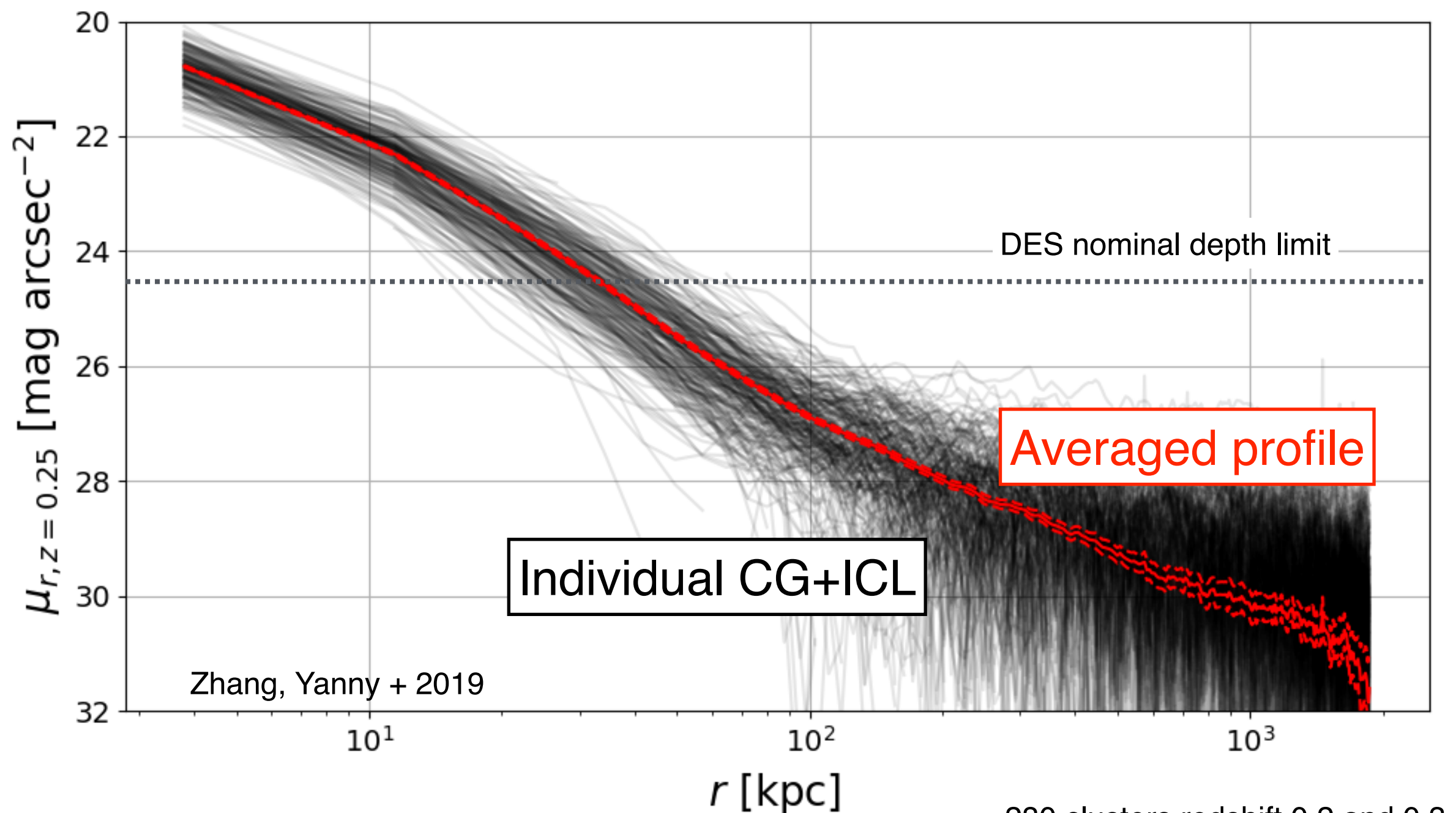


Step2: Additional background flux from “systematic” effects, such as nearby bright stars/galaxies, still need to be removed.



Step2: A second “background” evaluated through stacking “random points” in the DES footprint, is further subtracted.

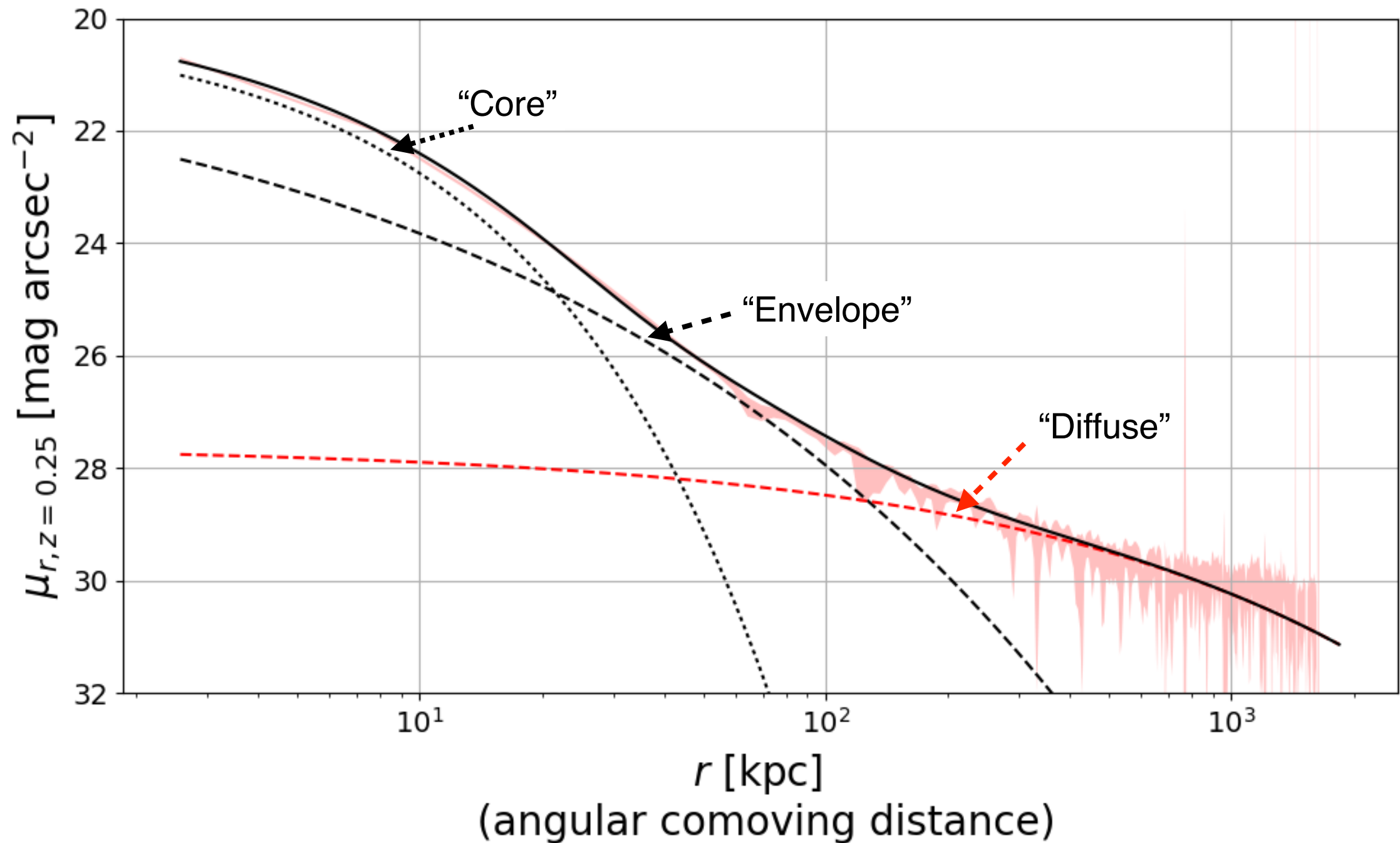
Thus, we were able to measure ICL out to ~ 1 Mpc from the cluster center and at a surface brightness level of 30 mag/arcsec².



~ 280 clusters redshift 0.2 and 0.3.
Based on DES Y3 imaging data.

Diffuse light can to be modeled by a composite of three **Sersic profiles**.

$$I(R) = I_e \exp \left\{ -b_n \left[\left(\frac{R}{R_e} \right)^{1/n} - 1 \right] \right\}$$



Summary and ongoing work

- “Stacking” helps detecting ICL out to ~ 1 Mpc at redshift 0.2 to 0.3.

See more details (sky background, PSF, cluster galaxy, ICL test) in:

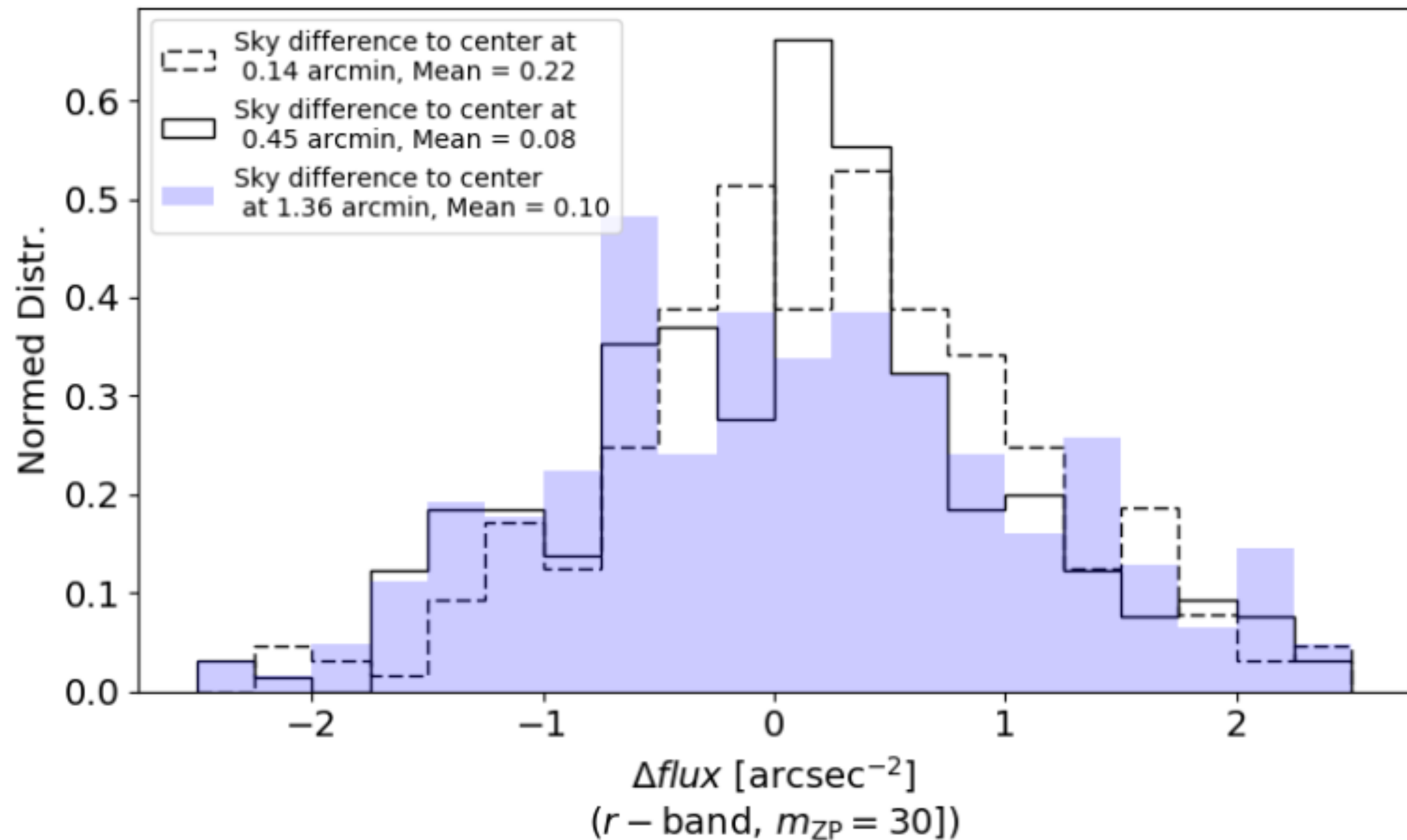
- ICL at redshift 0.2 to 0.3 — Zhang, Yanny + 2019, arXiv: 1812.04004
- ICL photo-z bias on lensing — Gruen, Zhang + 2018, arXiv:1809.04599
- Comparison of ICL to lensing — Sampaio-Santos, Zhang+ 2020, arXiv: 2005.12275
- Diffuse light profile of LRGs — Leung, Zhang+ 2020, arXiv: 2005.13467
- [Paper data release \(or scan QR code\)](#)

Ongoing work:

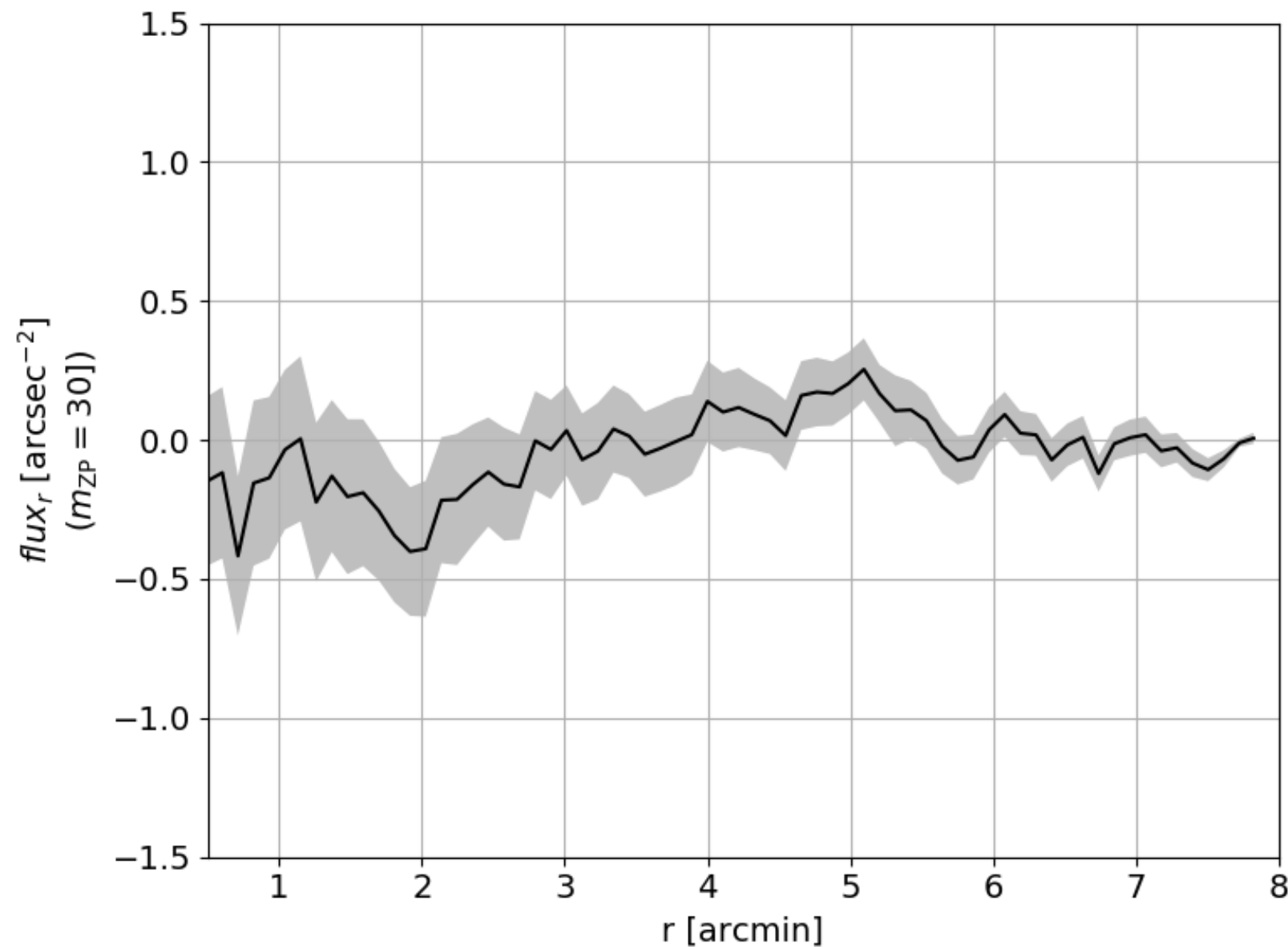
- ICL redshift evolution.



Sky background is estimated over the whole FOV, $\sim 3 \text{ deg}^2$ (Bernstein+2017), and appears well-behaved inside clusters.

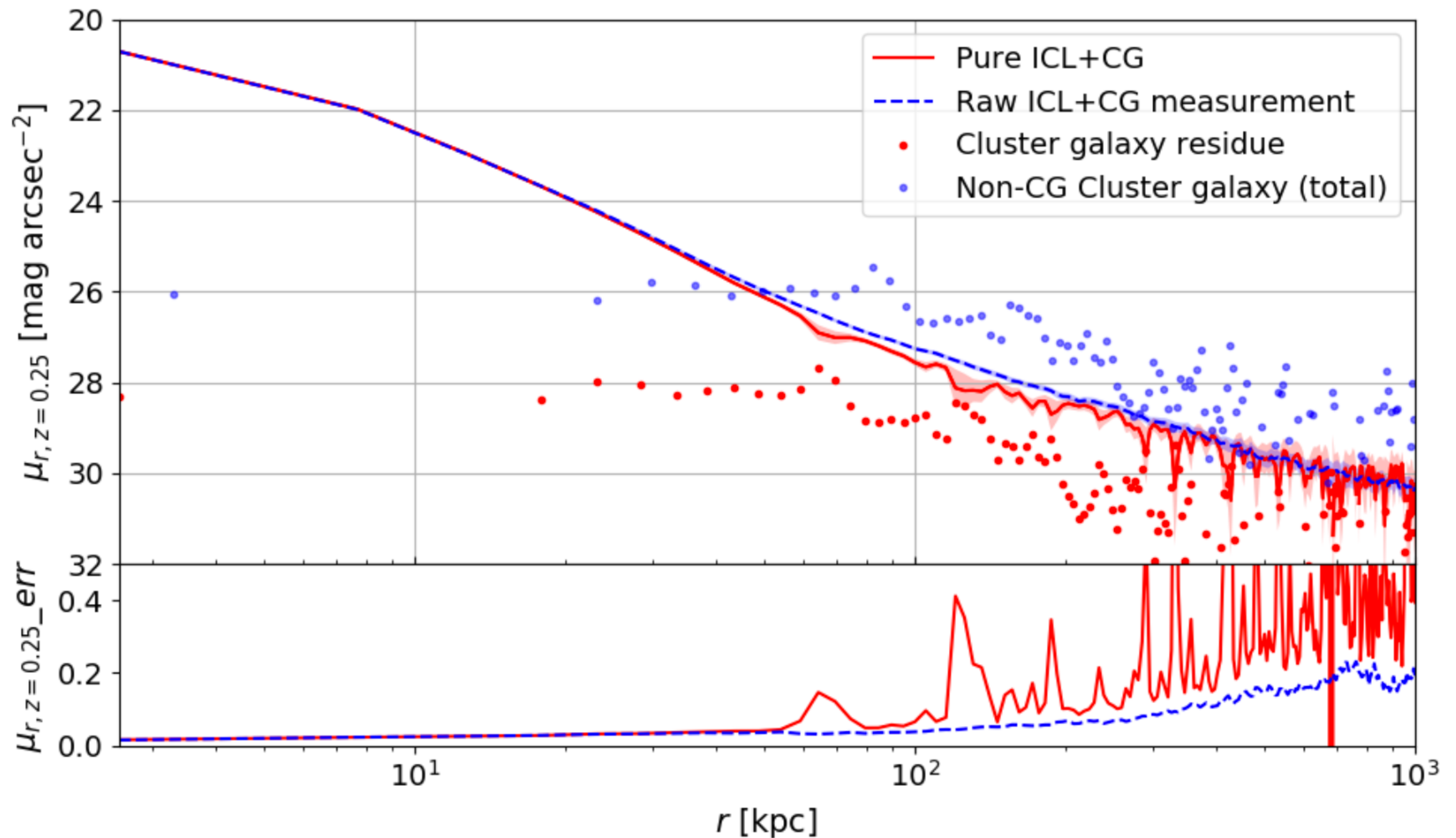


The measurement procedure is tested by “stacking random points”.



Stacking random points shows that the measurements are (relatively) bias-free.

Residual light of cluster member galaxies makes up a small component ($\sim 10\%$) of the diffuse light measurement.



Yes, diffuse light is an important component of galaxy cluster.

diffuse light makes up $44 \pm 17\%$
of the total cluster stellar luminosity within 1 Mpc.

