

## **Pushing the limits of source** detection tools towards LSB light

Nushkia Chamba<sup>1</sup>, Caroline Haigh<sup>2</sup>, Michael Wilkinson<sup>2</sup>, Aku Venhola<sup>3</sup> & Reynier Peletier<sup>2</sup>

LSST PCW, 10-14 August 2020

<sup>1</sup>Instituto de Astrofisica de Canarias, Tenerife, Spain <sup>2</sup>University of Groningen, Groningen, The Netherlands <sup>3</sup>University of Oulu, Oulu, Finland





22

Gobierno de Canarias

Consejería de Economía, Conocimiento y Empleo





Unión Europea Fondo Europeo de desarrollo Regional nera de hacer Europa





This project has received funding from the European Union's Horizon 2020 research and innovation programme under the Marie Sklodowska-Curie grant agreement No 721463



### **Everybody knows Source Extractor (SE)** Bertin & Arnouts (1996), but it has its limits



#### SDSS DR7 image, SE default settings

#### Teeninga, Moschini, Trager & Wilkinson (2016)





### Many other tools exist Can they be automatically optimised to detect LSB light?

# Source Extractor (SE) Bertin & Arnouts (1996) General purpose

### Profound (PF)

- •Robotham et al. (2018)
- General purpose

### **NoiseChisel (NC)**

- •Akhlaghi & Ishikawa (2015)
- Faint object specialised

### Max-Tree Objects (MT)

- •Teeninga et al. (2016)
- Faint object specialised

## **Basic characteristics**

- Measure background
- Threshold image w.r.t background
- Locate sources
- Catalogue and measure properties

	SE	PF	NC	MT
core method	nested thresholds	watershed	watershed	max-tree
initial threshold	$\lambda\sigma$	$\lambda\sigma$	percentile	0
nested objects		-	-	+
# thresholds	discrete	NA	NA	$\infty$
detection by statistical test				-
parallel	+	?	+	
# parameters	12	8	25	2

### Source finding using Trees **Max-Tree Objects (MT)**





#### SDSS DR7 image, MT, 2 relevant parameters

#### Teeninga, Moschini, Trager & Wilkinson (2016)



### A comparison of detection tools **Overview**

- In this work: SE, NC, PF and MTO
- arcsec<sup>2</sup>)]
- Automatic parameter optimisation
- Four different quality measures
- Tests on real images (FDS, IAC Stripe 82, Hubble Ultra Deep Field)

• Simulated deep data [Fornax Deep Survey,  $\mu_{lim} \sim 30$  mag/arcsec<sup>2</sup> (3 $\sigma$ ; 100

## Ground truth for faint light



#### Simulated FDS image



#### Ground truth at $0.1\sigma$

### **Evaluation** Quality criteria

F1 score: Combines precision (purity) and recall (completeness) in pure detection task

$$F_1 = \frac{\text{precision} \times \text{recall}}{\text{precision} + \text{recall}}$$

**Area score:** Optimizes segmentation quality, combining under-merging error (UM) and over-merging error (OM)

Area-score = 
$$1 - \sqrt{OM^2 + U}$$

**Combined score A:** 
$$\sqrt{\text{Area-score}^2 + F_1^2}$$

Combined score B:  $\sqrt[3]{(1 - OM)(1 - UM)F_1}$ 

 $UM^2$ 

### Evaluation **Parameter Optimisation**

- Ten simulated images are used
- Each of the settings is tested on the remaining 9 images

Bayesian optimisation is performed on each image for each quality measure

## **Results - Summary**





## **Real images - Two IAC Stripe 82 examples**



#### Galactic cirrus

 $\mu_{g,lim} = 29.1 \text{ mag/arcsec}^2 (3\sigma, 100 \text{ arcsec}^2)$ http://research.iac.es/proyecto/stripe82/



#### Tidal streams, bright sources



### **Results - Galactic cirrus SExtractor**



#### **Optimised for F-score**



### **Results - Galactic cirrus ProFound**



#### Optimised for F-score



### **Results - Galactic cirrus** NoiseChisel



#### Optimised for F-score





### **Results - Galactic cirrus** Max-Tree Objects



#### **Optimised for F-score**







### Tidal streams, bright sources

PF



Area









#### **Combined A**







### **Combined B**







### PF (in R): Too slow

## HUDF





NC

SE





PF

### UDF 423 Credit: NASA

Area









#### **Combined A**







#### **Combined B**







### PF (in R): Too slow

### Background values Talk to me for details!

- Mean background value of simulated image is zero
- Each algorithm has its internal estimator (can be imrpov)
- Both PF and SE consistently overestimated the background: O(10<sup>-1</sup> $\sigma$ )
- MT underestimated the value: area score O(-10<sup>-1</sup> $\sigma$ ) and F-score O(-10<sup>-2} $\sigma$ )</sup>
- NC showed the strongest performance:  $O(\pm 10^{-3}\sigma)$

### **Concluding remarks** How can these results help you?

- Robust, optimised parameters for detection algorithms. SCARLET?
- Evaluation: MT overall most stable and consistent performance (C. Haigh et al. re-submitted to A&A

	MTObjec
Optimised parameters	2
Language	Python/C
Clean edges of detected objects	(H)
Detects galaxy close to star (Stripe 82)	$\checkmark$
Detects cirrus (Stripe 82)	$\checkmark$
Isolates spiral substructures (HUDF)	$\checkmark$

Be aware (beware) of each algorithms limits and failures: do you care about nested objects or only faint outskirts? de-blending?

cts	NoiseChisel	ProFound	SExtractor
	20	8	6
С	С	R	С
	$\checkmark$	$\checkmark$	Sometimes
	Fragmented		Fragmented
	$\checkmark$	-	Sometimes





(b) A fragmented galaxy.

### Max Tree Objects: Concept **Component Trees**

Based on decomposition of image into connected components



Teeninga, Moschini, Trager & Wilkinson (2016)