

### Sensor Characterization and ISR



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#### Sensor Characterization and ISR

Created by John Swinbank, last modified by Andrés Alejandro Plazas Malagón on Apr 04, 2019

Goals: To Create a list/table of all effects related sensor characterization and ISR, and, for each one, understanding what the camera team is doing, what the DM team is doing, and, if there are any differences in the algorithms and/or implementations, understanding why.

Sources of information:

- Instrument Signature Removal and Calibration Production Products (input for ISR):
  - Calibration Products Production: cp\_pipe (generates calibration products): https://github.com/lsst/cp\_pipe
    - Chapter 4 ("Calibration Products Production", CPP, by Merlin) of LDM-151 ("Data Management Science Pipelines Design"): https://docushare.lsst.org/docushare/dsweb/Get/LDM-151
      - Inputs to CPP
        - Bias frames, gain values, linearity, darks, crosstalk, defect map, saturation levels, broadband flats, monochromatic flats, collimated beam projector data, filter transmission, atmospheric characterization.
      - Outputs from CPP == Inputs to the AP/DRP Pipelines (ISR pipelines, in particular):
        - Master bias, master darks, master linearity, master fringe frames, master gain values, master defects, saturation levels crosstalk, master impure broadband flats, master impure monochromatic flats, master pure monochromatic flats, master photo flats, master low-resolution narrow-band flats, pixel sizes, brighter-fatter coefficients, CTE measurement, filter transmission, ghost catalog, spectral standards, spectrophotometric standards, astrometric standards.
    - "LSST Plans for Calibrated Photometry": https://github.com/lsst-dm/calibration/blob/master/calibration.pdf
    - DMTN-101: "Verifying LSST Calibration Data Products" (https://dmtn-101.lsst.io/)
  - Instrument Signature Removal, ip\_isr (applies calibration products): https://github.com/lsst/ip\_isr
    - "Inputs to primary method of IsrTask, run(), are a raw exposure to be corrected and the calibration data products. The raw input is a single chip sized mosaic of all amps including overscans and other non-science pixels."
      - (from 'run()' in isrTask.py) ccdExposure, bias, linearizer (linearizing functor), dark, flat, defects (list of defects), fringes, bfKernel, camera (camera geometry, WCS?), opticsTransmission, filterTransmission, sensorTransmission, atmosphereTransmission, crosstalkSources (dictionary, default: DECam inter-CCD)
    - Current order in isrTask.py (as of 2/13/2019; Chris Waters (Princeton)):
      - 1) Integer to float conversion
      - 2) Bad amplifier and SATURATED/SUSPECT pixel masking
      - 3) Overscan correction
      - 4) Optional crosstalk correction (doCrossTalk)
      - 5) CDD assembly (doAssemble)
      - 6) Bias correction (doBias)
      - 7) Variance calculation (variance image construction, doVariance)
      - 8) Linearization (doLinearize)



## Sensor Characterization and ISR



Robert Lupton's slides on physical motivations for ISR order: isr.pdf

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### **Instrumental Signature Removal**

**Robert Lupton, Princeton University** 

2018-12-19



### Instrument Signature Removal



# ip\_isr, <a href="https://github.com/lsst/ip\_isr">https://github.com/lsst/ip\_isr</a>

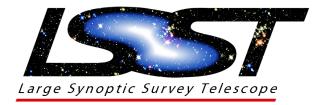
- 1) Integer to float conversion
- 2) Bad amplifier and SATURATED/SUSPECT pixel masking
- 3) Overscan correction
- 4) Optional crosstalk correction
- 5) CDD assembly
- 6) Bias correction
- 7) Variance calculation (variance image construction)
- 8) Linearization
- 9) Crosstalk
- 10) Saturation trail widening (HSC)
- 11) Brighter-Fatter correction
- 12) Dark
- 13) Optional fringe correction
- 14) Straylight
- 15) Flat/gain normalization
- 16) Defect masking and interpolation
- 17) Fringe
- 18) Bad pixel/camera specific masking
- 19) Attach vignette model (HSC), transmission curves, and distortion model

- 1) Integer to float conversion
- 2) Bad amplifier and SATURATED/SUSPECT pixel masking
- 3) Overscan correction
- 4) CCD Assembly
- 5) Bias correction
- 6) Crosstalk correction
- 7) Linearization
- 8) Charge transfer correction (not currently implemented)
- 9) Dark
- 10) Saturation trail widening (HSC)
- 11) Brighter-Fatter correction
- 12) Variance calculation
- 13) Straylight
- 14) Flat/gain normalization
- 15) Defect masking and interpolation
- 16) Fringe
- 17) Bad pixel/camera specific masking





- Algorithms: What could be added? What is needed? What should be prioritized?
  - Examples:
    - BFE:
      - Craig: tests on correction of BFE (> 90%)
      - Pierre^2: covs, models in "The Shape of the PTC in CCDs"
    - CTI: not currently implemented in ISR
      - High, low flux models: Adam
      - Thoughts by Pierre (DESC meeting at Berkeley)
    - Focal Plane Electro-Optical testing
      - Comparisons to what is in DM
      - Should any algorithm or analysis be implemented in the stack?
      - Eric's code
    - On-sky data:
      - HSC, DEcam, comCam, auxTel



# Thank you!

